Current numerical and physical limitations of the EAGLE and APOSTLE simulations. Some paths for the future.

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www.eaglesim.org -> Public Galaxy Catalogue

# EAGLE in a nutshell

EAGLE: Evolution and Assembly of GaLaxies and their Environments

The evolution of intergalactic gas. Colour encodes temperature

z = 14.0 t = 0.3 Gyr L = 25.0 cMpc

Visualisation by Jim Geach & Rob Crain

Schaye+2015 (1407.7040), Crain+2015 (1501.01311), Schaller+2015 (1509.05056)

# EAGLE in a nutshell

Free parameters in the model were calibrated to reproduce the z=0 GSMF and relation between galaxy masses and sizes.



 $\rightarrow$  Looks like we have a plausible mass-loading function (by construction !)



# A sizeable problem for cosmology



4.5M CPU hours (final run) + 40M hours (calibration) + 2M hours (post-processing)

# **Baryons-DM** interaction



- Halo mass functions change between DMONLY and HYDRO (>10% level) !
- If (big if) one could measure the HMF, and would use DMONLY as reference point, one would infer modified gravity at the many-sigma level !

See also Cusworth+2014 on the same topic.

### **Baryons-DM** interaction



Hellwing, Schaller et al. 2016

Total matter power-spectrum different (between n-body and EAGLE) at >1% level. Not a trivial effect:

- not easy to "fit" out of a DMONLY simulation.

#### Baryons-DM offset ?



Schaller et al. 2015

## Is getting the GSMF good enough ?



# Is getting the GSMF good enough ?

- There seem to be more to it.
- Different Mass-loading functions that give the ~ same GSMF lead to different results.
- Need more constrains to discriminate the Mass-loading functions.



Onset of quenching at too large a mass  $\rightarrow$  change AGN ?



HI not great at low masses (Rob's talk). Improved SPH is important here.



Mass – Metallicity too shallow ? Lack of mixing ? Resolution problem ?



Discs are clearly too thick  $\rightarrow$  EoS problem ? Softening problem ?

#### Equation of State



#### Equation of State



#### **Technical issues**

On 2048 cores, wallclock = 549.3 hours to redshift 0.00, timestep = 2.43731e+06



# Conclusions

- We have a reasonable "Mass loading function" (by construction !).
- Nevertheless, HI masses, metallicities and disk heights are not greatly reproduced.
- Higher resolution helps for the first two.
- Will need improved model (different EoS, different softening ? ).
- Calibration of the future runs will be very expensive and better scaling codes are required.